

THE CANADIAN MINERALOGIST

Journal of the Mineralogical Association of Canada

Volume 12

May, 1974

Part 5

Canadian Mineralogist
Vol. 12, pp. 293-298 (1974)

CAYSICHITE, A NEW SILICO-CARBONATE OF YTTRIUM AND CALCIUM

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ABSTRACT

Caysichite, $(Y,Ca)_2Si_2O_{10}(CO_3)_3 \cdot 4H_2O$, occurs in granite pegmatite at the abandoned Evans-Lou feldspar mine, 22 miles north of Ottawa. The mineral is normally present as a coating on fractures or as incrustations up to 2 mm thick with a divergent columnar structure. Terminated crystals are rare. Caysichite is colourless to white, more rarely yellow or green. $H = 4\frac{1}{2}$, $D(\text{obs}) = 3.03$ and $D(\text{calc}) = 3.029 \text{ g/cm}^3$. For the colourless variety, $a = 1.586$, $\beta = 1.614$, $\gamma = 1.621$, $2V_x = 53^\circ$, $X = b$, $Y = a$, $Z = c$, with positive elongation; yellow crystals have somewhat higher refractive indices and larger optic angle. The mineral is orthorhombic: $Ccm2_1$ or $Ccmm$ with $a = 13.282$, $b = 13.925$, $c = 9.727 \text{ \AA}$ (crystal elongation); $Z = 4$. Strongest lines ($CuK\alpha$) are 6.93 (020,111) (100), 4.38 (130) (60), 4.22 (310) (60), 3.48 (040) (60), 3.32 (400) (90). IR absorptions occur at $3700\text{-}2600 \text{ cm}^{-1}$ (H_2O), $1700\text{-}1240 \text{ cm}^{-1}$ (CO_3) and $1200\text{-}900 \text{ cm}^{-1}$ (SiO_4). On heating, dehydration is followed by two distinct stages of decarbonation. The DTA product (1120°C) has an apatite structure. Caysichite may have been derived from hellandite through low temperature solution and precipitation.

INTRODUCTION

Yttrium and lanthanide minerals are ubiquitous, though minor, constituents of Precambrian granite pegmatites that were once worked

for feldspar and quartz in eastern Ontario and western Quebec. The minerals were generally looked upon as a nuisance by the operators and were discarded on the dumps, where they are turned up periodically by collectors. One such rare earth pegmatite is the Evans-Lou feldspar mine, some 22 miles north of Ottawa. This mine was quarried for feldspar and quartz from 1932 to 1956 and collections from the dumps have since yielded several interesting mineral species (Hogarth 1972) including a new yttrium calcium silico-carbonate, described in this paper. The mineral was first observed on the mine dump in May, 1969, as a white powder, coating perthite, and as it could not be identified by x -ray diffraction methods, the mineral was designated UN-11 of the unidentified Evans-Lou minerals. In August of the same year yellowish microcrystals from cavities in quartz specimens from the dump were x -rayed and designated UN-16 (Miles *et al.* 1971). Later, the x -ray patterns of UN-16 and UN-11 were found to be similar, differing only in intensities.

In the summer of 1970, the quarry was partly drained of water to permit the examination and collection of samples *in situ*. After the draining, UN-16 was observed as faintly yellowish crusts on quartz along a 2-foot length in the hanging (west) wall of the pegmatite and as white radiating crystals, within a 3-foot radius, at the north

end of the quarry. These locations are 15 to 20 feet below the normal water level of the flooded quarry.

UN-16 has been named caysichite for its principal cations: Ca, Y, Si, C, H. The mineral and the name, syllabized cay-si-chite and pronounced KĀ.SICHAIT, have been approved by the Commission on New Minerals and Mineral Names, I.M.A. The type material has been deposited with the Systematic Reference Series (No. 10402) of the National Mineral Collection, Geological Survey of Canada, Ottawa.

OCCURRENCE

At the Evans-Lou mine, caysichite occurs in granite pegmatite. It is especially common in quartz-rich areas of the giant perthite zone which surrounds the quartz core (Miles *et al.* 1971) but has not been found in the core itself. The mineral coats fractures and encrusts or completely fills cavities in quartz, perthite and hellandite, and is apparently of supergene origin. It is most closely associated with hellandite but is also found with the other rare earth-bearing minerals fergusonite, kainosite (cenosite), tengerite, lokkaite and yttrian thorogummite; less commonly with yttrian spessartine, xerotime, wakefieldite and synchisite-(Y). Further details of its association are given by Hogarth (1972).

Most commonly, caysichite lines cavities as a dull white pulverulent coating or as a cream stain. More rarely it is found as thin, brittle incrustations with a parallel to slightly divergent columnar structure and with a reniform surface somewhat resembling prehnite (Fig. 1). In this form it is coherent, though easily separable into 'splinters' up to 2 mm long. Additionally, the mineral occurs as radiating groups and terminated crystals, and, at one locality, as stalactites up to one cm in length.

PHYSICAL AND MORPHOLOGICAL PROPERTIES

Colourless, crystalline crusts of caysichite on samples from the hanging wall of the pegmatite (location EL-344, Hogarth 1972) were the most amenable to clean separation, and splinters, hand-picked from these samples under the binocular microscope, provided the material for most studies. The hardness is $4\frac{1}{2}$ and the VHN is 551. The density, as determined by heavy liquids, is 3.03 g/cm^3 . Specimens are non-fluorescent in long and short wave ultraviolet radiation but show a faint green cathodoluminescence under electron bombardment. In hand specimen the mineral is colourless, white to pale yellow,

and rarely greenish. It has a vitreous lustre and white streak, and effervesces slowly in cold, dilute HCl. The rarely-observed caysichite crystals are prismatic, elongated parallel to *c* and terminated by {001} (Fig. 2). A prism and a dome



FIG. 1. Caysichite incrustations on quartz, Evans-Lou mine, Quebec.

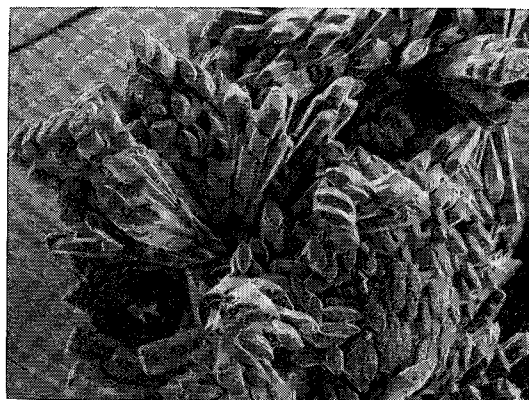


FIG. 2. Scanning electron photomicrograph of crystals of caysichite showing their prismatic habit and {001} termination, and also a poorly developed dome on some crystals. Width of crystal is 0.05 mm.

can be observed under high magnification but their Miller indices could not be determined because of poor signals on the optical goniometer.

The optical data (Table 1) were obtained at room temperature on the spindle stage using a sodium vapour lamp illuminator. All Cargille immersion liquids used were checked by an Abbe refractometer. The mineral is biaxial negative with parallel extinction and positive elongation. No dispersion was noted in any of the grains examined. In conventional immersion mounts, caysichite fragments give nearly centered acute bisectrix figures suggesting {010} cleavage.

TABLE 1. OPTICAL DATA FOR CAYSICHITE

Colourless crystal	Yellowish crystal
$\alpha = 1.586 \pm 0.004$	$\alpha = 1.589 \pm 0.004$
$\beta = 1.614 \pm 0.001$	$\beta = 1.616 \pm 0.001$
$\gamma = 1.621 \pm 0.001$	$\gamma = 1.626 \pm 0.001$
$2V_x = 53^\circ(\text{meas.}); 54^\circ 30'(\text{calc.})$	$2V_x = 72^\circ 45'(\text{calc.})$
$X = b, Y = a, Z = c$	non-pleochroic

X-RAY CRYSTALLOGRAPHY

Crystals suitable for Weissenberg and precession studies were obtained only with considerable difficulty because of the sub-parallel growth of many individuals. The best single-crystal photographs, obtained from what is presumably a minute cleavage fragment, indicate that caysichite is orthorhombic with $a = 13.30$, $b = 13.95$, and $c = 9.74\text{\AA}$. The observed systematic extinctions are consistent with those required by space groups $Ccm2_1$ and $Ccmm$.

X-ray powder diffraction data for the colourless caysichite are given in Table 2. With the single-crystal photographs as a guide, the powder pattern is readily indexed on the basis of the orthorhombic cell. The cell parameters derived from the single crystal studies were then refined by a least-squares method using the indexed powder data. The refined values are $a = 13.282(3)$, $b = 13.925(3)$ and $c = 9.724(4)\text{\AA}$.

The powder pattern of the yellow variety is identical to that of the colourless crystals except for very minor differences in d -values and for the relative intensities of a few lines.

CHEMICAL COMPOSITION

Compositional data for caysichite are presented in Table 3. Water was determined on carefully hand-picked samples by a modified Penfield method, and carbon dioxide by both titrimetry and by total weight loss minus water

TABLE 2. X-RAY POWDER DIFFRACTION PATTERN OF CAYSICHITE (COLOURLESS)

hkl	I/I_0	d_{obs}	d_{calc}	hkl	I/I_0	d_{obs}	d_{calc}
001			9.727	153	20	2.085	2.086
110	5	9.62	9.661	621	5	2.061	2.062
020			6.962	513	5	2.033	2.033
111	100	6.93	6.838	602	5	2.016	2.015
002	40	4.87	4.864	044	<5	1.994	1.994
130	60	4.38	4.382	404	5	1.965	1.962
310	60	4.22	4.219	234	10	1.937	1.937
022	10	3.985	3.987	622	10	1.937	1.935
311	10	3.866	3.871	550	5	1.920	1.922
040	60	3.481	3.481	460	10	1.904	1.902
400	90	3.321	3.320	710	30	1.879	1.880
132	5	3.257	3.255	533	30	1.879	1.879
330	30	3.201	3.204	263	10	1.814	1.815
401	10	3.146	3.142	370	10	1.814	1.814
240	30	3.077	3.083	514	10	1.778	1.779
113	30	3.077	3.072	462	10	1.778	1.772
420	<5	2.997	2.997	623	10	1.778	1.768
241	30	2.939	2.939	730	30	1.754	1.756
042	<5	2.829	2.831	712	30	1.754	1.754
402	30	2.733	2.742	372	<5	1.701	1.700
150	30	2.733	2.726	064	<5	1.680	1.679
332	30	2.677	2.675	553	20	1.653	1.654
133	30	2.608	2.606	732	30	1.652	1.652
313	10	2.574	2.571	082	5	1.637	1.639
511	5	2.517	2.520	604	5	1.637	1.637
004	5	2.434	2.432	660	10	1.600	1.602
440	<5	2.405	2.403	155	5	1.582	1.580
152	5	2.376	2.378	661	5	1.582	1.584
060	40	2.323	2.321	802	5	1.573	1.571
530	5	2.305	2.306	190	10	1.538	1.537
204	5	2.286	2.284	662	10	1.520	1.521
531	5	2.240	2.243	752	5	1.491	1.492
243	5	2.240	2.234	390	5	1.461	1.461
600	30	2.213	2.214				
601	30	2.156	2.158				
442	10	2.156	2.154				
134	10	2.125	2.126				
352	10	2.120	2.121				
620	20	2.108	2.110				
314	20	2.108	2.107				

plus other lines

114.6 mm camera, $\text{CuK}\alpha$ radiation ($\lambda=1.5418\text{\AA}$), Si standard, visual intensities

(loss on ignition minus H_2O). All other quantitative data were obtained by electron probe analysis on ten optically clear splinters of colourless caysichite, less than 1 mm long and 0.2-0.3 mm

TABLE 3. CHEMICAL COMPOSITION OF CAYSICHITE

Wt. %	Atoms per 19 Oxygens	
	actual	ideal
CaO	10.04	Ca 1.473
Y_2O_3	28.18	Y 2.054
La_2O_3	0.08	La 0.004
CeO_2	0.24	Ce 0.011
Pr_2O_3	0.09	Pr 0.004
Nd_2O_3	0.30	Nd 0.015
Sm_2O_3	0.20	Sm 0.009
Eu_2O_3	0.09	Eu 0.004
Gd_2O_3	0.51	Gd 0.023
Tb_2O_3	0.16	Tb 0.007
Dy_2O_3	1.16	Dy 0.051
Ho_2O_3	0.66	Ho 0.029
Er_2O_3	1.84	Er 0.079
Tm_2O_3	0.21	Tm 0.009
Yb_2O_3	2.12	Yb 0.089
Lu_2O_3	0.42	Lu 0.017
SiO_2	28.84	Si 3.950
Al_2O_3	0.58	Al 0.094
CO_2	15.7*	C 2.937
H_2O	8.6**	H_2O 3.932
Total	100.02	

3.878 4.000

4.044 4.000

*total wt. loss on ignition of 17 mg of sample, less H_2O ; 14.1% by titrimetry on 25 mg.

**modified Penfield method.

Additional minor elements by mass spectroscopy (wt. %): B 0.05, F 0.04, P 0.01, Mg 0.02, K 0.02, Sc 0.01, Cr 0.01, Mn 0.01, Ni 0.03, Cu 0.01, Sr 0.02, Sn 0.01, Ba 0.01, Pb 0.02.

wide. Energy dispersive analysis spectra of the splinters showed Ca, Y and Si to be the principal elements and lanthanides to be present in minor amounts. Quantitative determinations for these elements were carried out by conventional methods with a Materials Analysis Company electron microprobe using synthetic yttrium-iron garnet, kainosite (cenosite of Pouliot *et al.* 1964) and synthetic rare earth element glasses (Drake & Weill 1972) as standards. To avoid sample decomposition a beam diameter of 10 microns was employed, with a specimen current of 0.03 microamperes. The data were processed by using a revised and updated version of the computer program by Rucklidge (1967). The data in Table 3 are the average of at least 10 spot analyses for each element. The compositional range of 50 points in 5 grains for Y and Ca was: Y_2O_3 27.4-28.8% and CaO 10.3-9.8%, with the elements varying antipathetically. The elements Mg, Ba, Ti, Sr, Na, K, S, P, Cl and F were sought by electron microprobe analysis, and each found to be less than 0.05%. Fe, as high as 0.2%, was detected at a few points in some grains. Minor elements detected by mass spectrometry are noted at the bottom of the table.

The analysis, recalculated on the basis of 19 oxygen atoms, gives the following formula: $Y_{2.05}Ca_{1.47}RE_{3.35}Si_{3.95}Al_{0.06}O_{10.18}(CO_3)_{2.94} \cdot 3.93H_2O$, or ideally: $(Y,Ca)_2Si_4O_{10}(CO_3)_3 \cdot 4H_2O$. Assum-

ing four formula units per cell the calculated density is 3.029 g/cm^3 , which is in excellent agreement with the measured value.

Following identification of the yellowish variety of caysichite by optical and x-ray methods, a pale greenish yellow, cone-shaped grain, $1.2 \times 1.1 \text{ mm}$ and tapering to 0.3 mm , was mounted for analysis. In polished section, the greater part of the grain consists of interpenetrating acicular crystals of normal caysichite, which take a good polish, and which have rhomb-shaped cross-sections. Locally, the grain is rimmed by a veneer, up to 50 microns thick, of a calcium- and yttrium-rich mineral that is probably synchisite-(Y). Towards the margin of the grain, interstitial areas between the well-crystallized caysichite are filled with very finely crystalline material which has a poorer polish. This material is chemically similar to normal caysichite, except for a higher ratio of Y to Ca, with 32% Y_2O_3 and 7% CaO, and a small increase to 0.5% for Fe. The variations noted in the optical and x-ray data for the colourless and yellowish caysichite may therefore be due directly to variations in the Y/Ca ratios of the mineral.

THERMAL STUDIES

Thermal studies were carried out on four different hand-picked samples. Thermogravimetric analysis in air (Fig. 3) shows three stages of weight loss: (1) $290 - 580^\circ\text{C}$ (8.3%) probably representing dehydration, (2) $580 - 895^\circ\text{C}$ (10.6%), thought to represent a first stage decarbonation, (3) $895 - 1120^\circ$ (5.3%) thought to represent a second stage decarbonation. The total weight loss was 24.3% and the end product gave an x-ray diffraction pattern of an apatite structure. The specimen had intumesced but not melted at the end-temperature (1120°C). These data can be correlated with the differential thermal analysis curve (Fig. 4) which shows two prominent endotherms at 431°C and 801°C and minor peaks from $935 - 1100^\circ\text{C}$.

In static heating studies, a sample heated in air for 15 minutes at 350°C showed a weight loss of 2.84%; for 15 minutes at 550°C , a loss of 9.37% (cumulative); and for 50 minutes at 1100°C , a loss of 24.85% (cumulative).

INFRARED STUDIES

The infrared spectra of caysichite are presented in Figure 5. The strong, broad band in the region $3700 - 2600 \text{ cm}^{-1}$ is attributable to hydrogen-bonded water. This is consistent with the TGA and DTA observations that caysichite de-

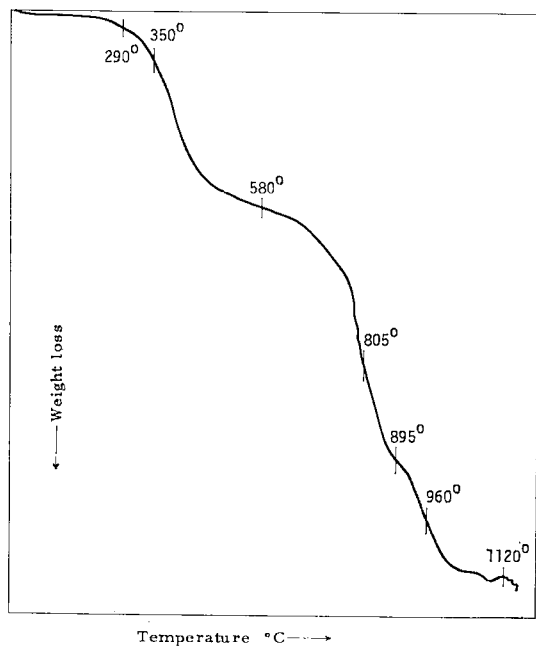


FIG. 3. Thermogravimetric curve of caysichite (heating rate = $12^\circ\text{C}/\text{min}$, in air).

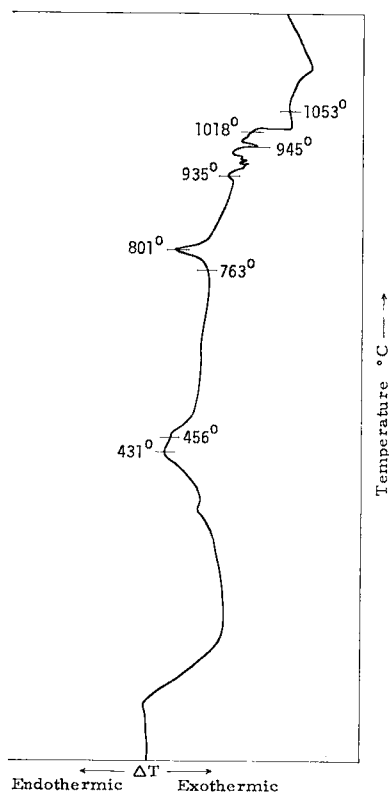


FIG. 4. Differential thermal analysis curve of caysichite, (heating rate = $10^{\circ}\text{C}/\text{min}$, in CO_2).

hydrates at a relatively high temperature (431 – 435°C). The strong bands in the region 1700 – 1240 cm^{-1} are due to the vibrations of the CO_3 group. The multiplicity and poor resolution of bands in this region may be due to the presence of two distinct CO_3 groups in the structure, as suggested by the two-stage decarbonation observed in the TGA study. The strong bands in the region 1200 – 900 cm^{-1} are attributed to Si-O stretching. The assignments of the minor bands in the spectra are: 3020 – 2800 cm^{-1} (overtone of ν_3 of CO_3); 1720 cm^{-1} (adsorbed water); 830 cm^{-1} (ν_2 of CO_3); 790 and 740 cm^{-1} (Si-Si stretching); 690 , 630 and 620 cm^{-1} (ν_4 of CO_3).

ORIGIN

The close spatial relationship of caysichite and hellandite suggests that caysichite is secondary after hellandite. However, the distribution of lanthanides in Evans-Lou hellandite, caysichite and fergusonite (Fig. 6) shows a closer

chemical relationship between caysichite and fergusonite. The absence of a gadolinium peak in caysichite may be ascribed to crystal-chemical control, the structure of caysichite being more selective than that of hellandite for the smaller and therefore the heavier members of the lanthanide group.

The crustiform, reniform and stalactitic nature of caysichite, as a partial or complete filling of cavities or open fractures, strongly suggests that it was formed at low or moderate temperatures. Solution and transportation may have taken place in mildly acidic carbonated groundwaters, with precipitation occurring when the pH was increased to a value approaching neutrality.

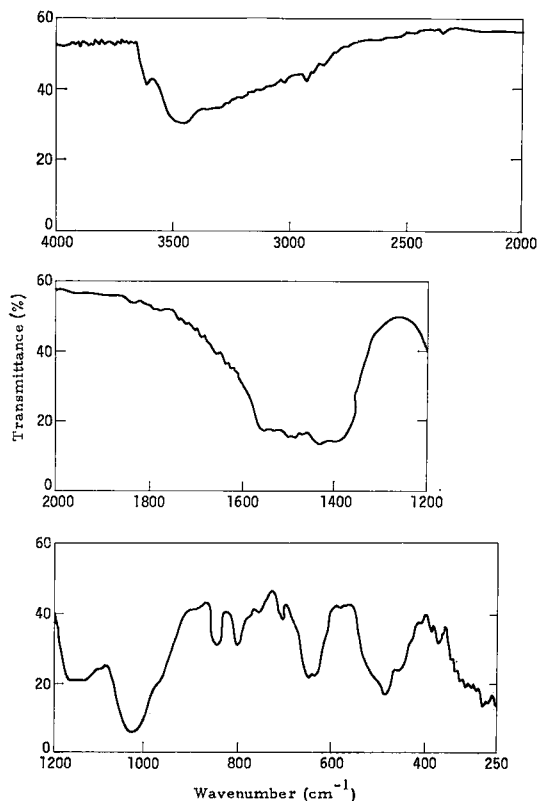


FIG. 5. Infrared spectra of caysichite.

ACKNOWLEDGEMENTS

We wish to express our thanks to Mr. Leo MacDonnell of the International Minerals and Chemical Corporation (Canada) Limited for granting us permission to visit the property and collect specimens. Grateful acknowledgement is

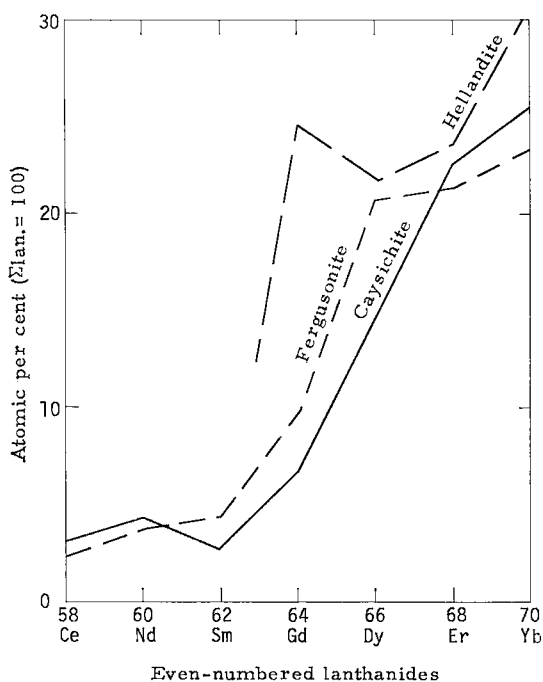


FIG. 6. Distribution of lanthanides in caysichite, hellandite and fergusonite, Evans-Lou mine, Quebec.

also made to D. S. Russell, National Research Council, for the mass spectrometric analysis, to R. Lake, Mines Branch, for DTA and TGA data and, at the Geological Survey of Canada, to J. G. Sen Gupta for the chemical analyses for CO_2 and H_2O , and to C. R. McLeod for the Vickers microhardness tests. The study was supported by a grant from the National Research Council to D. D. Hogarth, which he gratefully acknowledges.

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